Exact results in AdS/CFT from localization

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Based on work with Fernando Alday, Daniel Farquet, Martin Fluder, Carolina Gregory Jakob Lorenzen, Dario Martelli, and Paul Richmond

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Field theory:

- In the last few years it has been appreciated that one can put general (Euclidean) supersymmetric gauge theories on curved backgrounds, preserving supersymmetry.
- In such a theory the VEV of any BPS operator localizes

$$
\langle \mathcal{O}_{\text{BPS}} \rangle = \int_{\text{all fields}} e^{-S} \mathcal{O}_{\text{BPS}} \cdot \left(\text{one-loop determinant} \right).
$$

$$
= \int_{\mathcal{Q}-\text{invariant fields}} e^{-S} \mathcal{O}_{\text{BPS}} \cdot \left(\text{one-loop determinant} \right).
$$

A form of fixed point theorem: Q is a supercharge, generating a supersymmetry variation of the theory.

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- For appropriate classes of theories and operators one can compute such quantities exactly in field theory, on an arbitrary background M_d .
- Applications include non-perturbative tests of various conjectured dualities.

In particular, if the field theory on (conformally) flat space has an AdS dual, we may try to compare these computations to gravity.

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Gravity:

- There are large classes of supersymmetric gauge theories that, in a suitable large N limit, are conjectured to be described by the supergravity limit of string/M-theory.
- Typically described by a (warped) product $AdS_{d+1} \times Y$, where different choices of internal space Y correspond to different gauge theories, and $N =$ flux quantum number.
- We must then solve a supergravity filling problem in Euclidean quantum gravity: find the (least action) solution on some M_{d+1} such that $\partial M_{d+1} = M_d$.

I will summarize results for $d = 3$ and $d = 5$.

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One can put an arbitrary $\mathcal{N} = 2$ supersymmetric gauge theory in $d = 3$ dimensions on a (Euclidean) curved background following [Festuccia-Seiberg]: couple the theory to $d = 3$ supergravity, and take a rigid limit in which $m_{\text{Planck}} \rightarrow \infty$ [Closset-Dumitrescu-Festuccia-Komargodski].

As well as the background metric on M_3 , there are two background vector fields A and V, and a scalar function h, together with Killing spinor χ satisfying

$$
(\nabla_{\mu} - iA_{\mu})\chi = -\frac{i}{2}h\gamma_{\mu}\chi - iV_{\mu}\chi - \frac{1}{2}\epsilon_{\mu\nu\rho}V^{\nu}\gamma^{\rho}\chi.
$$

Of central importance for us is the Killing vector

$$
K = \chi^{\dagger} \gamma^{\mu} \chi \partial_{\mu} = \partial_{\psi} .
$$

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The vector field **K** is nowhere zero, generating a foliation of M_3 which is transversely holomorphic. The metric is locally

$$
\mathrm{d} s_3^2 = \Omega(z,\bar{z})^2 (\mathrm{d}\psi + a)^2 + c(z,\bar{z})^2 \mathrm{d} z \mathrm{d}\bar{z}.
$$

where **z** is a complex coordinate.

Essentially the background is parametrized by an arbitrary choice of the functions $\Omega(z, \bar{z})$, $c(z, \bar{z})$, and local one-form $a = a(z, \bar{z})dz + c.c.$, and imposing the Killing spinor equation then fixes everything else in terms of these.

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If all the orbits of K close then M_3 is the total space of a $U(1)$ orbibundle over an orbifold Riemann surface Σ (a Seifert fibred 3-manifold).

On the other hand, if at least one orbit is open then M_3 necessarily admits a $U(1) \times U(1)$ isometry, and we may write

$$
\mathsf{K} \;=\; \partial_{\psi} \;=\; \mathsf{b}_1 \partial_{\varphi_1} + \mathsf{b}_2 \partial_{\varphi_2} \;,
$$

where $\mathbf{b}_1, \mathbf{b}_2 \neq \mathbf{0}$ can be thought of as parametrizing a choice of **K**.

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General $\mathcal{N} = 2$ supersymmetric gauge theory in $d = 3$ dimensions:

- Vector multiplet $({\mathscr{A}}, \sigma, \lambda, D)$ in the adjoint of the gauge group **G**, for which we may write a Chern-Simons, as well as Yang-Mills, action.
- Matter chiral multiplet (ϕ, ψ, F) in a representation $\mathcal R$ of **G**, with superpotential.

The localization computation for this general set-up is in our paper [1307.6848]. One first determines the Q -invariant field configurations, and then computes the one-loop determinants around these.

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In the vector multiplet we find the localization equations for $\mathsf{M}_{3}\cong\mathsf{S}^{3}$ imply

$$
\mathscr{A}~=~\mathbf{0}~, \qquad \Omega\sigma~=~\sigma_0~=~{\rm constant}~, \qquad \mathsf{D}=-\frac{\mathsf{h}}{\Omega}\sigma_0~.
$$

The matter multiplet is trivial: all fields localize to zero.

The classical action for $\mathsf{M}_3\cong\mathsf{S}^3$, evaluated on the localization locus, is given entirely by the Chern-Simons action:

$$
\mathsf{S}_{\mathrm{CS}}\;=\;-\frac{\mathrm{i}\mathsf{k}}{2\pi}\mathrm{Tr}(\sigma_0^2)\int_{\mathsf{M}_3}\frac{\mathsf{h}}{\Omega^2}\sqrt{\mathsf{det}\,\mathsf{g}}\,\mathrm{d}^3\mathsf{x}\;=\;\frac{\mathrm{i}\pi\mathsf{k}}{|\mathsf{b}_1\mathsf{b}_2|}\mathrm{Tr}(\sigma_0^2)\;.
$$

Most of the work is in computing the one-loop determinants.

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The final result for the partition function is

$$
Z = \int d\sigma_0 e^{-\frac{i\pi k}{|b_1b_2|}\text{Tr}\,\sigma_0^2} \prod_{\alpha \in \Delta_+} 4 \sinh \frac{\pi \sigma_0 \alpha}{|b_1|} \sinh \frac{\pi \sigma_0 \alpha}{|b_2|}
$$

$$
\cdot \prod_{\rho} s_{\beta} \left[\frac{i(\beta + \beta^{-1})}{2} (1 - R) - \frac{\rho(\sigma_0)}{\sqrt{|b_1b_2|}} \right].
$$

Here we have defined $\beta=\sqrt{|\mathbf{b}_1/\mathbf{b}_2|}$, ρ denote weights in a weight space decomposition of the representation $\mathcal R$ for the matter fields, **R** is their R-charge, and $s_\beta(z)$ denotes the double sine function.

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It is also straightforward to insert BPS operators, for example the Wilson loop

$$
W = Tr_{\mathcal{R}} \left[\mathcal{P} \exp \int_{\gamma} ds (i \mathscr{A}_{\mu} \dot{x}^{\mu} + \sigma |\dot{x}|) \right],
$$

where $\mathsf{x}^\mu(\mathsf{s})$ parametrizes with worldline $\gamma =$ orbit of $\mathsf{K}% (\mathsf{A})$ is \mathcal{Q}_{γ} -invariant.

 \langle **W** \rangle is then computed by inserting $\text{Tr}_{\mathcal{R}}\text{e}^{2\pi \ell \sigma_0}$ into the localized partition function, where $2\pi \ell =$ length of Reeb orbit (e.g. at the "pole" where $\partial_{\varphi_1} = 0$, $\ell = 1/|b_2|$ [Farquet-JFS].

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For comparison with AdS/CFT we should focus on field theories that in (conformally) flat space have an AdS gravity dual.

There are huge classes of these, described by Chern-Simons-quiver gauge theories, with $\mathsf{U}(\mathsf{N})^{\mathsf{p}}$ gauge groups, *e.g.* the maximally supersymmetric case is the ABJM theory, living on N M2-branes in flat space.

The gravity duals are M-theory backgrounds of the form $AdS_4 \times Y_7$, with N units of $*G_4$ through the internal space Y_7 , and arise as e.g. near-horizon limits of N M2-branes at Calabi-Yau four-fold singularities [Martelli-JFS, many other authors].

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The large N limit of the matrix model partition function was computed in [Martelli-Passias-JFS], using a saddle point method of [Herzog-Klebanov-Pufu-Tesileanu].

This involves the asymptotic expansion of the double sine function, and an ansatz for the saddle point eigenvalue distribution for σ_0 .

The final results are extremely simple:

$$
\log Z = \frac{(|b_1|+|b_2|)^2}{4|b_1b_2|} \cdot \log Z_{\text{round }S^3} ,
$$

$$
\log \langle W \rangle = \frac{1}{2} \ell (|b_1|+|b_2|) \cdot \log \langle W \rangle_{\text{round }S^3} .
$$

In particular, the dependence on the background geometry factorizes from the dependence on the choice of gauge theory.

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In [1404.0268] and [1406.2493] we have reproduced these formulas from a dual Euclidean quantum gravity calculation, for a very general class of solutions.

We work in $\mathcal{N} = 2$ gauged supergravity in four dimensions. This is Einstein-Maxwell theory, with a graviphoton \mathcal{A} , and we use the fact that any supersymmetric solution of this theory on M_4 uplifts to a supersymmetric solution of M-theory on $M_4 \times Y_7$ [Gauntlett-Varela].

The Killing spinor equation takes the form

$$
\left[\nabla_{\mu}-\mathrm{i}\mathcal{A}_{\mu}+\frac{1}{2}\Gamma_{\mu}+\frac{\mathrm{i}}{4}\mathcal{F}_{\nu\rho}\Gamma^{\nu\rho}\Gamma_{\mu}\right]\epsilon\;=\;0\;.
$$

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The local form of Euclidean supersymmetric solutions to this theory was studied by [Dunajski-Gutowski-Sabra-Tod].

In particular, there is a class of *self-dual* solutions in which $*_4\mathcal{F} = -\mathcal{F}$ is anti-self-dual, and the four-metric is then Einstein with anti-self-dual Weyl tensor.

We also have a Killing vector

$$
K = i\epsilon^{\dagger} \Gamma^{\mu} \Gamma_5 \epsilon \partial_{\mu} = \partial_{\psi} .
$$

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Self-dual Einstein metrics with a Killing vector have a rich geometric structure. They are (locally) conformal to a scalar-flat Kähler metric, with the metric determined entirely by a solution to the Toda equation:

$$
\mathrm{d} s_4^2 \; = \; \frac{1}{y^2} \mathrm{d} s_{\mathrm{Kahler}}^2 \; = \; \frac{1}{y^2} \Big[V^{-1} (\mathrm{d} \psi + \phi)^2 + V (\mathrm{d} y^2 + 4 \mathrm{e}^{\mathsf{w}} \mathrm{d} z \mathrm{d} \bar{z}) \Big] \; .
$$

where ${\sf V}={\sf 1}-\frac12{\sf y}\partial_{\sf y}{\sf w}$, the expression for ${\rm d}\phi$ is known (but complicated), and

$$
\partial_z \partial_{\bar z} w + \partial_y^2 e^w = 0.
$$

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The conformal boundary is at $y = 0$, and one can show that the structure induced on the conformal boundary is precisely the three-dimensional background geometry of [Closset-Dumitrescu-Festuccia-Komargodski].

In particular

$$
\epsilon = y^{-1/2} \left[(1 + \Gamma_0 + \tfrac{1}{4} y w_{(1)} \Gamma_0) \begin{pmatrix} \chi \\ 0 \end{pmatrix} + \mathcal{O}(y^2) \right] ,
$$

where χ is a three-dimensional spinor satisfying the Killing spinor equation we saw earlier, and we expand ${\sf w}({\sf y},{\sf z},\bar{\sf z})={\sf w}_{(0)}({\sf z},\bar{\sf z})+{\sf y}{\sf w}_{(1)}({\sf z},\bar{\sf z})+\mathcal{O}({\sf y}^2).$

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Suppose we have such a solution. The holographic free energy is

$$
-\log Z_{gravity} \; = \; S_{Einstein-Maxwell} + S_{Gibbons-Hawking} + S_{counterterms} \; .
$$

The individual terms certainly depend on the detailed solution. For example

$$
\frac{1}{16\pi G_N}\int_{B_4} F^2\sqrt{\det g}\,d^4x = -\frac{\pi(|b_1+b_2|)^2}{8G_N|b_1b_2|} + \frac{1}{256\pi G_N}\int_{M_3} \left(3w_{(1)}^3 + 4w_{(1)}w_{(2)}\right)\sqrt{\det g_3}\,d^3x.
$$

Here we have assumed the topology $\mathsf{M}_3\cong \mathsf{S}^3$ and $\mathsf{M}_4\cong \mathsf{B}_4$.

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However, the final result is

$$
-\log Z_{\text{gravity}} = \frac{(|b_1|+|b_2|)^2}{4|b_1b_2|} \cdot \frac{\pi}{2G_N} \ ,
$$

agreeing with the field theory computation!

The Wilson loop in the fundamental representation maps to a supersymmetric M2-brane, wrapping a calibrated copy of the M-theory circle [Farquet-JFS], and with a minimal surface $\Sigma \subset \mathsf{B}_4$ with $\partial \Sigma = \gamma = 0$ orbit of Reeb vector **K**.

 $log\langle W \rangle_{gravity}$ is identified with minus the regularized action of the M2-brane, and in $[1406.2493]$ we showed this reproduces the large N field theory result.

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We now change focus to $d = 5$. [Imamura] has defined five-dimensional supersymmetric gauge theories on the $SU(3) \times U(1)$ -invariant squashed five-sphere background

$$
ds_5^2 = \frac{1}{s^2}(d\tau + C)^2 + ds_{\mathbb{CP}^2}^2
$$

where $\frac{1}{2}$ dC = ω = Kähler form for the Fubini-Study metric on \mathbb{CP}^2 . Here s = squashing parameter, with $s = 1$ the round five-sphere.

There is also a background R-symmetry gauge field

$$
A^R = \frac{1}{s^2}(1 + Q\sqrt{1 - s^2})\sqrt{1 - s^2}(d\tau + C) ,
$$

where $U(1)_R \subset SU(2)_R$ and $Q = 1$, $Q = -3$ give rise to 3/4 BPS and 1/4 BPS solutions, respectively.

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The *perturbative* partition function again localizes onto an integral over the constant mode σ_0 of the scalar in the vector multiplet, and the final formula involves triple sine functions.

A particular class of five-dimensional gauge theories, with gauge group USp(2N) and arising from a $D4-D8$ system, is expected to have a large N description in terms of massive type IIA supergravity [Ferrara-Kehagias-Partouche-Zaffaroni], [Brandhuber-Oz].

In \int Jafferis-Puful the large N limit of the partition function of these theories on the round sphere was computed and successfully compared to the entanglement entropy of the dual warped $\mathsf{AdS}_6\times\mathsf{S}^4$ supergravity solution.

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In $[1405.7194]$ we computed the large N limit of the USp(2N) gauge theories on the squashed five-sphere, finding the free energy

$$
\log Z = \frac{(|b_1| + |b_2| + |b_3|)^3}{27|b_1b_2b_3|} \cdot \log Z_{\text{round }S^5} \ ,
$$

where

$$
\begin{cases}\n\mathbf{b}_1 = \mathbf{b}_2 = \mathbf{b}_3 & 1/4 \text{ BPS} \\
\mathbf{b}_1 = -1 - \sqrt{1 - s^2}, \ \mathbf{b}_2 = \mathbf{b}_3 = 1 - \sqrt{1 - s^2} & 3/4 \text{ BPS}\n\end{cases}
$$

There is again a supersymmetric Killing vector bilinear K, and embedding $\mathsf{S}^5\subset\mathbb{R}^2\oplus\mathbb{R}^2\oplus\mathbb{R}^2$, this is $\mathsf{K}=\mathsf{b}_1\partial_{\varphi_1}+\mathsf{b}_2\partial_{\varphi_2}+\mathsf{b}_3\partial_{\varphi_3}.$

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We also computed the large N limit of BPS Wilson loops. If the worldline wraps the $\textsf{S}_\textsf{i}^1\subset \textsf{S}^5$ at the origin of two copies of \mathbb{R}^2 , then we find

$$
\log \langle W \,\rangle \; = \; \frac{|b_1|+|b_2|+|b_3|}{3|b_i|} \cdot \log \langle W \,\rangle_{\mathrm{round\,}S^5} \; .
$$

We have reproduced these formulae from a dual supergravity computation.

We work in six-dimensional Romans $F(4)$ gauged supergravity, which is a consistent truncation of massive IIA supergravity on ${\sf S}^4$ [Cvetic-Lu-Pope]. As well as the metric, there is a scalar X , two-form potential B , one-form potential A , and an $\mathsf{SO}(3)\sim \mathsf{SU}(2)$ R-symmetry gauge field $\mathsf{A}_\mathsf{I},\,\mathsf{I}=1,2,3$.

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The one-form **A** is a Stueckelberg field, which may be set to $A = 0$ by a gauge transformation. The B-field then becomes massive, and the Euclidean action is

$$
\begin{array}{lll} S_{\rm bulk} & = & - \displaystyle \frac{1}{16 \pi G_N} \int_{M_6} \left[R*1 - 4 X^{-2} {\rm d} X \wedge * {\rm d} X \right. \\ & & \left. - \left(\frac{2}{9} X^{-6} - \frac{8}{3} X^{-2} - 2 X^2 \right) * 1 - \frac{1}{2} X^{-2} \left(\frac{4}{9} B \wedge * B + F_1 \wedge * F_1 \right) \right. \\ & & \left. - \frac{1}{2} X^4 H \wedge * H - i B \wedge \left(\frac{2}{27} B \wedge B + \frac{1}{2} F_1 \wedge F_1 \right) \right] \, . \end{array}
$$

Notice the cubic Chern-Simons coupling for **B**. Its curvature is $H = dB$.

A solution to the corresponding equations of motion is supersymmetric provided the Killing spinor equation and dilatino equation hold.

The squashed five-sphere background has $SU(3) \times U(1)$ symmetry, and one expects this to be preserved by the bulk filling. This leads to the ansatz

$$
ds_6^2 = \alpha^2(r)dr^2 + \gamma^2(r)(d\tau + C)^2 + \beta^2(r)ds_{\mathbb{CP}^2}^2,
$$

\n
$$
B = p(r)dr \wedge (d\tau + C) + \frac{1}{2}q(r)dC,
$$

\n
$$
A_1 = f_1(r)(d\tau + C),
$$

together with $X = X(r)$.

We have constructed smooth, supersymmetric, asymptotically locally Euclidean AdS solutions with the topology $M_6 \cong B_6$, with conformal boundary the squashed five-sphere backgrounds of [Imamura]. These may be given as expansions around the conformal boundary $r = \infty$, and/or as expansions in the squashing parameter s.

Reparametrization invariance allows us to set $\beta(\mathsf{r})=3\sqrt{6}\mathsf{r}^2-1/2$ √ 2 to its AdS₆ value, and an SO(3) rotation sets $f_3(r) = f(r)$, $f_1(r) = f_2(r) = 0$.

For example, for the 3/4 BPS solution the first few terms in the expansion around $r = \infty$ are

$$
\alpha(r) = \frac{3}{\sqrt{2}}r + \frac{8+s^2}{36\sqrt{2s^2}}\frac{1}{r^3} + \cdots,
$$

\n
$$
\gamma(r) = \frac{3\sqrt{3}}{s}r + \frac{-16+7s^2}{12\sqrt{3s^3}}\frac{1}{r} - \frac{-1280+1120s^2+241s^4}{2592\sqrt{3s^5}}\frac{1}{r^3} + \cdots,
$$

\n
$$
X(r) = 1 + \frac{1-s^2-3\sqrt{1-s^2}}{54s^2}\frac{1}{r^2} + \frac{s^2\sqrt{1-s^2}\kappa}{12(1-s^2+\sqrt{1-s^2})}\frac{1}{r^3} + \cdots,
$$

\n
$$
\rho(r) = -\frac{i\sqrt{\frac{2}{3}}\left(s^2+3\sqrt{1-s^2}-1\right)}{s^3}\frac{1}{r^2} + \cdots,
$$

\n
$$
q(r) = -\frac{3i\left(\sqrt{6}\sqrt{1-s^2}\right)}{s}r + \frac{\sqrt{\frac{2}{3}}i\sqrt{1-s^2}\left(5s^2+9\sqrt{1-s^2}-5\right)}{3s^3}\frac{1}{r} + \cdots,
$$

\n
$$
f(r) = \frac{1-s^2+\sqrt{1-s^2}}{s^2} + \frac{2\left(-2+2s^2-(2+s^2)\sqrt{1-s^2}\right)}{9s^4}\frac{1}{r^2} + \frac{\kappa}{r^3} + \cdots.
$$

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The parameter κ is uniquely determined by requiring this to extend to a smooth solution on the ball $M_6 \cong B_6$. As an expansion in

$$
\delta = \sqrt{-1 + s^{-1}}
$$

this is

$$
\frac{3\sqrt{3}}{4}\kappa = \delta + \frac{\sqrt{2}}{3}\delta^2 + \frac{113}{36}\delta^3 + \frac{25}{9\sqrt{2}}\delta^4 + \frac{1127}{288}\delta^5 + \frac{35}{9\sqrt{2}}\delta^6 + \dots
$$

Similar results hold in the 1/4 BPS case, except here we find a two-parameter family of solutions, leading to a new supersymmetric squashing of S^5 . In particular this includes a one-parameter subfamily of 1/2 BPS solutions.

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As in four dimensions the regularized action is

$$
-\log Z_{gravity} = S_{bulk} + S_{Gibbons-Hawking} + S_{ct} .
$$

However, unlike in four dimensions the counterterms S_{ct} had not been computed.

This is a straightforward, but very long, computation. In particular the B-field is both massive and has a cubic Chern-Simons interaction, which leads to a much more complicated analysis than for more standard fields.

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$$
\begin{array}{lcl} S_{ct} & = & \displaystyle \frac{1}{8\pi G_N} \int_{\partial M_{\tilde{D}}} \left\{ \Big[\frac{4\sqrt{2}}{3} + \frac{1}{2\sqrt{2}} R(h) - \frac{1}{6\sqrt{2}}\|B\|_{h}^{2} + \frac{3}{4\sqrt{2}}R(h)_{ij}R(h)^{ij} - \frac{15}{64\sqrt{2}}R(h)^{2} - \frac{3}{4\sqrt{2}}\|F_{I}\|_{h}^{2} \right. \\ & & \left. + \displaystyle \frac{1}{12\sqrt{2}} {\rm Tr}_{h}B^{4} + \displaystyle \frac{5}{8\sqrt{2}}\|d\ast_{h}B + \displaystyle \frac{{\rm i}\sqrt{2}}{3}B\wedge B\|_{h}^{2} - \displaystyle \frac{1}{4\sqrt{2}}(B, {\rm d}\delta_{h}B + \displaystyle \frac{{\rm i}\sqrt{2}}{3} {\rm d}\ast_{h}B\wedge B)_{h} - \displaystyle \frac{1}{\sqrt{2}}\|{\rm d}B\|_{h}^{2} \right. \\ & & \left. + \displaystyle \frac{4\sqrt{2}}{3} (1-X)^{2} - \displaystyle \frac{1}{\sqrt{2}}\langle {\rm Ric}(h)\circ B, B\rangle_{h} + \displaystyle \frac{9}{32\sqrt{2}}R(h)\|B\|_{h}^{2} - \displaystyle \frac{13}{192\sqrt{2}}\|B\|_{h}^{4} \right] \sqrt{\rm det}\,h\, {\rm d}^{5}x \\ & & \left. - \displaystyle \frac{1}{4\sqrt{2}} B\wedge \left[{\rm d}\ast_{h} {\rm d}B + \displaystyle \frac{\sqrt{2}{\rm i}}{3}B\wedge \delta_{h}B - \displaystyle \frac{2}{9}B\wedge \ast_{h}(B\wedge B)\right] \right\} \, . \end{array}
$$

Here $\text{Ric}(\mathbf{h})_{ii} = \mathbf{R}(\mathbf{h})_{ii}$ denotes the Ricci tensor of the boundary metric \mathbf{h}_{ii} , with **R(h)** the Ricci scalar. The inner product of two **p**-forms ν_1 , ν_2 is defined by $\langle \nu_1, \nu_2 \rangle$ _h $\sqrt{\det h} \, d^5x = \nu_1 \wedge *_h \nu_2$, which then also defines the square norm via $\|\nu\|_{\mathsf{h}}^2 = \langle \nu, \nu \rangle_{\mathsf{h}}$. The adjoint δ_{h} of ${\rm d}$ with respect to h_{ij} acting on the two-form **B** is $\delta_h B = *_{h} d *_{h} B$, and we have also defined $\text{Tr}_{h} B^4 \equiv B_i^{\ j} B_j^{\ k} B_k^{\ l} B_i^{\ i}$. Finally, we have defined the **p**-form $(\sf{S} \circ \nu)_{i_1\cdots i_p} \equiv \sf{S}_{[i_1}{}^j\nu_{|j|i_2\cdots i_p]},$ where \sf{S}_{ij} is any symmetric 2-tensor, and ν is any p-form.

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Using this we may compute the holographic free energy. For example, for the 3/4 BPS solution we find

$$
S_{bulk} + S_{Gibbons-Hawking} + S_{ct} = -\frac{27\pi^2}{4G_N} \left(1 + \frac{8}{3} \delta^2 + \frac{16\sqrt{2}}{27} \delta^3 + \frac{68}{27} \delta^4 + \frac{28\sqrt{2}}{27} \delta^5 + \frac{32}{27} \delta^6 + \ldots \right) .
$$

This agrees with the field theory result. The BPS Wilson loop maps to a fundamental string in type **IIA**, at the "pole" of the internal S^4 [Assel-Estes-Yamazaki]. The renormalized string action is

$$
S_{\rm string} = \int_{\Sigma} \left[X^{-2} \sqrt{\det \gamma} \, d^2 x + iB \right] - \frac{3}{\sqrt{2}} \mathrm{length}(\partial \Sigma) ,
$$

and also agrees with the large N field theory results.

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 $\mathbf{A} \oplus \mathbf{B} \rightarrow \mathbf{A} \oplus \mathbf{B} \rightarrow \mathbf{A} \oplus \mathbf{B} \rightarrow \mathbf{B}$