Exact results in AdS/CFT from localization

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Field theory:

- In the last few years it has been appreciated that one can put general (Euclidean) supersymmetric gauge theories on curved backgrounds, preserving supersymmetry.
- In such a theory the VEV of any BPS operator localizes

$$\begin{array}{ll} \langle \, \mathcal{O}_{\mathsf{BPS}} \, \rangle & = & \int_{\mathsf{all fields}} \mathrm{e}^{-\mathsf{S}} \, \mathcal{O}_{\mathsf{BPS}} \\ & \stackrel{\mathrm{exactly}}{=} & \int_{\mathcal{Q}-\mathrm{invariant fields}} \mathrm{e}^{-\mathsf{S}} \, \mathcal{O}_{\mathsf{BPS}} \cdot \text{(one-loop determinant)} \, . \end{array}$$

A form of *fixed point theorem*: Q is a supercharge, generating a supersymmetry variation of the theory.

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- For appropriate classes of theories and operators one can compute such quantities *exactly* in field theory, on an *arbitrary* background **M**_d.
- Applications include non-perturbative tests of various conjectured dualities.

In particular, if the field theory on (conformally) flat space has an AdS dual, we may try to compare these computations to gravity.

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Gravity:

- There are large classes of supersymmetric gauge theories that, in a suitable large **N** limit, are conjectured to be described by the supergravity limit of string/M-theory.
- Typically described by a (warped) product $AdS_{d+1} \times Y$, where different choices of internal space Y correspond to different gauge theories, and N = flux quantum number.
- We must then solve a supergravity filling problem in Euclidean quantum gravity: find the (least action) solution on some M_{d+1} such that $\partial M_{d+1} = M_d$.

I will summarize results for d = 3 and d = 5.

One can put an arbitrary $\mathcal{N} = 2$ supersymmetric gauge theory in $\mathbf{d} = 3$ dimensions on a (Euclidean) curved background following [Festuccia-Seiberg]: couple the theory to $\mathbf{d} = 3$ supergravity, and take a rigid limit in which $\mathbf{m_{Planck}} \rightarrow \infty$ [Closset-Dumitrescu-Festuccia-Komargodski].

As well as the background metric on M_3 , there are two background vector fields **A** and **V**, and a scalar function **h**, together with Killing spinor χ satisfying

$$(
abla_{\mu} - \mathrm{i} \mathsf{A}_{\mu})\chi = -\frac{\mathrm{i}}{2} \mathsf{h} \gamma_{\mu} \chi - \mathrm{i} \mathsf{V}_{\mu} \chi - \frac{1}{2} \epsilon_{\mu\nu\rho} \mathsf{V}^{\nu} \gamma^{\rho} \chi \; .$$

Of central importance for us is the Killing vector

$$\mathsf{K} \;=\; \chi^{\dagger} \gamma^{\mu} \chi \partial_{\mu} \;=\; \partial_{\psi} \;.$$

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The vector field K is nowhere zero, generating a foliation of M_3 which is transversely holomorphic. The metric is locally

$$\mathrm{d} s_3^2 \;=\; \, \varOmega(\mathsf{z},\bar{\mathsf{z}})^2 (\mathrm{d} \psi + \mathsf{a})^2 + \mathsf{c}(\mathsf{z},\bar{\mathsf{z}})^2 \mathrm{d} \mathsf{z} \mathrm{d} \bar{\mathsf{z}} \;.$$

where \mathbf{z} is a complex coordinate.

Essentially the background is parametrized by an arbitrary choice of the functions $\Omega(z, \overline{z}), c(z, \overline{z})$, and local one-form $a = a(z, \overline{z})dz + c.c.$, and imposing the Killing spinor equation then fixes everything else in terms of these.

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If all the orbits of K close then M_3 is the total space of a U(1) orbibundle over an orbifold Riemann surface Σ (a Seifert fibred 3-manifold).

On the other hand, if at least one orbit is open then M_3 necessarily admits a $\mathsf{U}(1)\times\mathsf{U}(1)$ isometry, and we may write

$$\mathsf{K} \;=\; \partial_\psi \;=\; \mathsf{b}_1 \partial_{\varphi_1} + \mathsf{b}_2 \partial_{\varphi_2} \;,$$

where $\mathbf{b}_1, \mathbf{b}_2 \neq \mathbf{0}$ can be thought of as parametrizing a choice of K.

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General $\mathcal{N} = 2$ supersymmetric gauge theory in $\mathbf{d} = \mathbf{3}$ dimensions:

- Vector multiplet (A, σ, λ, D) in the adjoint of the gauge group G, for which we may write a Chern-Simons, as well as Yang-Mills, action.
- Matter chiral multiplet (ϕ, ψ, F) in a representation \mathcal{R} of G, with superpotential.

The localization computation for this general set-up is in our paper [1307.6848]. One first determines the Q-invariant field configurations, and then computes the one-loop determinants around these.

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In the vector multiplet we find the localization equations for $\mathsf{M}_3\cong\mathsf{S}^3$ imply

$$\mathscr{A} = \mathbf{0} \ , \qquad \Omega \sigma = \sigma_0 = ext{constant} \ , \qquad \mathsf{D} = - rac{\mathsf{n}}{\Omega} \sigma_0 \ .$$

The matter multiplet is trivial: all fields localize to zero.

The classical action for $M_3\cong S^3,$ evaluated on the localization locus, is given entirely by the Chern-Simons action:

$$\mathsf{S}_{\mathrm{CS}} \;=\; -\frac{\mathrm{i} \mathsf{k}}{2\pi} \mathrm{Tr}(\sigma_0^2) \int_{\mathsf{M}_3} \frac{\mathsf{h}}{\varOmega^2} \sqrt{\det g} \,\mathrm{d}^3 \mathsf{x} \;=\; \frac{\mathrm{i} \pi \mathsf{k}}{|\mathsf{b}_1 \mathsf{b}_2|} \mathrm{Tr}(\sigma_0^2) \;.$$

Most of the work is in computing the one-loop determinants.

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The final result for the partition function is

$$\begin{aligned} \mathsf{Z} &= \int \mathrm{d}\sigma_0 \, \mathrm{e}^{-\frac{\mathrm{i}\pi \mathrm{k}}{|\mathbf{b}_1 \mathbf{b}_2|} \operatorname{Tr} \sigma_0^2} \prod_{\alpha \in \Delta_+} 4 \sinh \frac{\pi \sigma_0 \alpha}{|\mathbf{b}_1|} \sinh \frac{\pi \sigma_0 \alpha}{|\mathbf{b}_2|} \\ &\cdot \prod_{\rho} \mathsf{s}_{\beta} \left[\frac{\mathrm{i}(\beta + \beta^{-1})}{2} (1 - \mathsf{R}) - \frac{\rho(\sigma_0)}{\sqrt{|\mathbf{b}_1 \mathbf{b}_2|}} \right] \,. \end{aligned}$$

Here we have defined $\beta = \sqrt{|\mathbf{b}_1/\mathbf{b}_2|}$, ρ denote weights in a weight space decomposition of the representation \mathcal{R} for the matter fields, **R** is their R-charge, and $\mathbf{s}_{\beta}(\mathbf{z})$ denotes the double sine function.

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It is also straightforward to insert BPS operators, for example the Wilson loop

$$\mathsf{W} = \mathrm{Tr}_{\mathcal{R}}\left[\mathcal{P}\exp\int_{\gamma}\mathrm{d}\mathbf{s}(\mathrm{i}\mathscr{A}_{\mu}\dot{\mathsf{x}}^{\mu}+\sigma|\dot{\mathsf{x}}|)
ight]\,,$$

where $\mathbf{x}^{\mu}(\mathbf{s})$ parametrizes with worldline $\gamma = \text{orbit of } \mathbf{K}$, is \mathcal{Q} -invariant.

 $\langle \mathbf{W} \rangle$ is then computed by inserting $\mathrm{Tr}_{\mathcal{R}} \mathrm{e}^{2\pi\ell\sigma_0}$ into the localized partition function, where $2\pi\ell = \text{length}$ of Reeb orbit (*e.g.* at the "pole" where $\partial_{\varphi_1} = \mathbf{0}$, $\ell = 1/|\mathbf{b}_2|$) [Farquet-JFS].

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For comparison with AdS/CFT we should focus on field theories that in (conformally) flat space have an AdS gravity dual.

There are huge classes of these, described by Chern-Simons-quiver gauge theories, with $U(N)^p$ gauge groups, *e.g.* the maximally supersymmetric case is the ABJM theory, living on N M2-branes in flat space.

The gravity duals are M-theory backgrounds of the form $AdS_4 \times Y_7$, with N units of $*G_4$ through the internal space Y_7 , and arise as *e.g.* near-horizon limits of N M2-branes at Calabi-Yau four-fold singularities [Martelli-JFS, many other authors].

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The large **N** limit of the matrix model partition function was computed in [Martelli-Passias-JFS], using a saddle point method of [Herzog-Klebanov-Pufu-Tesileanu].

This involves the asymptotic expansion of the double sine function, and an ansatz for the saddle point eigenvalue distribution for σ_0 .

The final results are extremely simple:

$$\begin{split} \log \mathsf{Z} &=& \frac{\left(|\mathsf{b}_1|+|\mathsf{b}_2|\right)^2}{4|\mathsf{b}_1\mathsf{b}_2|} \cdot \log \mathsf{Z}_{\mathrm{round}\,\mathsf{S}^3} \;,\\ \log \langle \,\mathsf{W}\,\rangle &=& \frac{1}{2}\ell(|\mathsf{b}_1|+|\mathsf{b}_2|) \cdot \log \langle \,\mathsf{W}\,\rangle_{\mathrm{round}\,\mathsf{S}^3} \;. \end{split}$$

In particular, the dependence on the background geometry factorizes from the dependence on the choice of gauge theory.

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In [1404.0268] and [1406.2493] we have reproduced these formulas from a dual Euclidean quantum gravity calculation, for a very general class of solutions.

We work in $\mathcal{N} = 2$ gauged supergravity in four dimensions. This is Einstein-Maxwell theory, with a graviphoton \mathcal{A} , and we use the fact that any supersymmetric solution of this theory on M_4 uplifts to a supersymmetric solution of M-theory on $M_4 \times Y_7$ [Gauntlett-Varela].

The Killing spinor equation takes the form

$$\left[\nabla_{\mu} - \mathrm{i}\mathcal{A}_{\mu} + \frac{1}{2}\Gamma_{\mu} + \frac{\mathrm{i}}{4}\mathcal{F}_{\nu\rho}\Gamma^{\nu\rho}\Gamma_{\mu}\right]\epsilon = \mathbf{0}.$$

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The local form of Euclidean supersymmetric solutions to this theory was studied by [Dunajski-Gutowski-Sabra-Tod].

In particular, there is a class of *self-dual* solutions in which $*_4\mathcal{F} = -\mathcal{F}$ is anti-self-dual, and the four-metric is then Einstein with anti-self-dual Weyl tensor.

We also have a Killing vector

$$\mathsf{K} = \mathrm{i}\epsilon^{\dagger}\Gamma^{\mu}\Gamma_{5}\epsilon\partial_{\mu} = \partial_{\psi} .$$

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Self-dual Einstein metrics with a Killing vector have a rich geometric structure. They are (locally) conformal to a scalar-flat Kähler metric, with the metric determined entirely by a solution to the Toda equation:

$$ds_4^2 = \frac{1}{y^2} ds_{Kahler}^2 = \frac{1}{y^2} \Big[V^{-1} (d\psi + \phi)^2 + V (dy^2 + 4e^w dz d\bar{z}) \Big] .$$

where $V = 1 - \frac{1}{2}y\partial_y w$, the expression for $d\phi$ is known (but complicated), and

$$\partial_z \partial_{\bar{z}} w + \partial_y^2 e^w = 0$$

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The conformal boundary is at $\mathbf{y} = \mathbf{0}$, and one can show that the structure induced on the conformal boundary is precisely the three-dimensional background geometry of [Closset-Dumitrescu-Festuccia-Komargodski].

In particular

$$\epsilon \ = \ \mathbf{y}^{-1/2} \left[\left(\mathbf{1} + \varGamma_0 + \frac{1}{4} \mathbf{y} \mathbf{w}_{(1)} \varGamma_0 \right) \left(\begin{array}{c} \boldsymbol{\chi} \\ \mathbf{0} \end{array} \right) + \mathcal{O}(\mathbf{y}^2) \right] \ ,$$

where χ is a three-dimensional spinor satisfying the Killing spinor equation we saw earlier, and we expand $w(y, z, \overline{z}) = w_{(0)}(z, \overline{z}) + yw_{(1)}(z, \overline{z}) + \mathcal{O}(y^2)$.

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Suppose we have such a solution. The holographic free energy is

$$-\log Z_{gravity} = S_{Einstein-Maxwell} + S_{Gibbons-Hawking} + S_{counterterms}$$
 .

The individual terms certainly depend on the detailed solution. For example

$$\begin{split} \frac{1}{16\pi G_N} \int_{B_4} F^2 \sqrt{\det g} \, \mathrm{d}^4 x \; = \; - \frac{\pi (|b_1 + b_2|)^2}{8G_N |b_1 b_2|} \\ & + \frac{1}{256\pi G_N} \int_{M_3} \left(3 w^3_{(1)} + 4 w_{(1)} w_{(2)} \right) \sqrt{\det g_3} \, \mathrm{d}^3 x \; . \end{split}$$

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Here we have assumed the topology $\mathsf{M}_3\cong\mathsf{S}^3$ and $\mathsf{M}_4\cong\mathsf{B}_4.$

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However, the final result is

$$-\log Z_{\text{gravity}} = rac{(|\mathbf{b}_1| + |\mathbf{b}_2|)^2}{4|\mathbf{b}_1\mathbf{b}_2|} \cdot rac{\pi}{2\mathsf{G}_{\mathsf{N}}} \; ,$$

agreeing with the field theory computation!

The Wilson loop in the fundamental representation maps to a supersymmetric M2-brane, wrapping a calibrated copy of the M-theory circle [Farquet-JFS], and with a minimal surface $\Sigma \subset B_4$ with $\partial \Sigma = \gamma =$ orbit of Reeb vector K.

 $\log\langle W \rangle_{gravity}$ is identified with minus the regularized action of the M2-brane, and in [1406.2493] we showed this reproduces the large **N** field theory result.

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We now change focus to d = 5. [Imamura] has defined five-dimensional supersymmetric gauge theories on the $SU(3) \times U(1)$ -invariant squashed five-sphere background

$$ds_{5}^{2} = \frac{1}{s^{2}}(d\tau + C)^{2} + ds_{\mathbb{CP}^{2}}^{2}$$

where $\frac{1}{2}d\mathbf{C} = \omega = K$ ähler form for the Fubini-Study metric on \mathbb{CP}^2 . Here $\mathbf{s} =$ squashing parameter, with $\mathbf{s} = \mathbf{1}$ the round five-sphere.

There is also a background R-symmetry gauge field

$$\mathsf{A}^{\mathsf{R}} \; = \; \frac{1}{s^2} (1 + \mathsf{Q}\sqrt{1-s^2}) \sqrt{1-s^2} (\mathsf{d}\tau + \mathsf{C}) \; ,$$

where $U(1)_R \subset SU(2)_R$ and Q = 1, Q = -3 give rise to 3/4 BPS and 1/4 BPS solutions, respectively.

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The *perturbative* partition function again localizes onto an integral over the constant mode σ_0 of the scalar in the vector multiplet, and the final formula involves triple sine functions.

A particular class of five-dimensional gauge theories, with gauge group USp(2N) and arising from a D4-D8 system, is expected to have a large N description in terms of massive type IIA supergravity [Ferrara-Kehagias-Partouche-Zaffaroni], [Brandhuber-Oz].

In [Jafferis-Pufu] the large **N** limit of the partition function of these theories on the *round* sphere was computed and successfully compared to the entanglement entropy of the dual warped $AdS_6 \times S^4$ supergravity solution.

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In [1405.7194] we computed the large N limit of the USp(2N) gauge theories on the squashed five-sphere, finding the free energy

$$\log Z \; = \; \frac{(|b_1| + |b_2| + |b_3|)^3}{27|b_1b_2b_3|} \cdot \log Z_{\rm round \, S^5} \; , \label{eq:cond_state}$$

where

$$\begin{cases} b_1 = b_2 = b_3 & 1/4 \text{ BPS} \\ b_1 = -1 - \sqrt{1 - s^2} \text{ , } b_2 = b_3 = 1 - \sqrt{1 - s^2} & 3/4 \text{ BPS} \end{cases}$$

There is again a supersymmetric Killing vector bilinear K, and embedding $S^5 \subset \mathbb{R}^2 \oplus \mathbb{R}^2 \oplus \mathbb{R}^2$, this is $K = b_1 \partial_{\varphi_1} + b_2 \partial_{\varphi_2} + b_3 \partial_{\varphi_3}$.

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We also computed the large N limit of BPS Wilson loops. If the worldline wraps the $S_i^1 \subset S^5$ at the origin of two copies of \mathbb{R}^2 , then we find

$$\log \langle \, W \, \rangle \ = \ \frac{|b_1| + |b_2| + |b_3|}{3|b_i|} \cdot \log \langle \, W \, \rangle_{\mathrm{round}\, S^5} \ .$$

We have reproduced these formulae from a dual supergravity computation.

We work in six-dimensional Romans F(4) gauged supergravity, which is a consistent truncation of massive IIA supergravity on S^4 [Cvetic-Lu-Pope]. As well as the metric, there is a scalar X, two-form potential B, one-form potential A, and an $SO(3) \sim SU(2)$ R-symmetry gauge field A_I , I = 1, 2, 3.

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The one-form **A** is a Stueckelberg field, which may be set to $\mathbf{A} = \mathbf{0}$ by a gauge transformation. The **B**-field then becomes massive, and the Euclidean action is

$$\begin{split} S_{\rm bulk} &= -\frac{1}{16\pi G_N} \int_{M_6} \left[R*1 - 4X^{-2} {\rm d} X \wedge * {\rm d} X \right. \\ &\quad \left. - \left(\tfrac{2}{9} X^{-6} - \tfrac{8}{3} X^{-2} - 2X^2 \right) * 1 - \tfrac{1}{2} X^{-2} \left(\tfrac{4}{9} B \wedge * B + F_I \wedge * F_I \right) \right. \\ &\quad \left. - \tfrac{1}{2} X^4 H \wedge * H - {\rm i} B \wedge \left(\tfrac{2}{27} B \wedge B + \tfrac{1}{2} F_I \wedge F_I \right) \right] \,. \end{split}$$

Notice the cubic Chern-Simons coupling for **B**. Its curvature is H = dB.

A solution to the corresponding equations of motion is supersymmetric provided the Killing spinor equation and dilatino equation hold.

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The squashed five-sphere background has $SU(3) \times U(1)$ symmetry, and one expects this to be preserved by the bulk filling. This leads to the ansatz

$$\begin{split} \mathrm{d} \mathsf{s}_6^2 &= \alpha^2(\mathsf{r}) \mathrm{d} \mathsf{r}^2 + \gamma^2(\mathsf{r}) (\mathrm{d} \tau + \mathsf{C})^2 + \beta^2(\mathsf{r}) \mathrm{d} \mathsf{s}_{\mathbb{CP}^2}^2 \ , \\ \mathsf{B} &= \mathsf{p}(\mathsf{r}) \mathrm{d} \mathsf{r} \wedge (\mathrm{d} \tau + \mathsf{C}) + \frac{1}{2} \mathsf{q}(\mathsf{r}) \mathrm{d} \mathsf{C} \ , \\ \mathsf{A}_\mathsf{I} &= \mathsf{f}_\mathsf{I}(\mathsf{r}) (\mathrm{d} \tau + \mathsf{C}) \ , \end{split}$$

together with X = X(r).

We have constructed smooth, supersymmetric, asymptotically locally Euclidean AdS solutions with the topology $M_6\cong B_6$, with conformal boundary the squashed five-sphere backgrounds of [Imamura]. These may be given as expansions around the conformal boundary $r=\infty$, and/or as expansions in the squashing parameter s.

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Reparametrization invariance allows us to set $\beta(\mathbf{r}) = 3\sqrt{6r^2 - 1/\sqrt{2}}$ to its AdS₆ value, and an SO(3) rotation sets $f_3(\mathbf{r}) = f(\mathbf{r})$, $f_1(\mathbf{r}) = f_2(\mathbf{r}) = 0$.

For example, for the 3/4 BPS solution the first few terms in the expansion around $\mathbf{r}=\infty$ are

$$\begin{split} &\alpha(r) &= \ \frac{3}{\sqrt{2}}r + \frac{8+s^2}{36\sqrt{2}s^2}\frac{1}{r^3} + \dots, \\ &\gamma(r) &= \ \frac{3\sqrt{3}}{s}r + \frac{-16+7s^2}{12\sqrt{3}s^3}\frac{1}{r} - \frac{-1280+1120s^2+241s^4}{2592\sqrt{3}s^5}\frac{1}{r^3} + \dots, \\ &X(r) &= \ 1 + \frac{1-s^2-3\sqrt{1-s^2}}{54s^2}\frac{1}{r^2} + \frac{s^2\sqrt{1-s^2}\kappa}{12\left(1-s^2+\sqrt{1-s^2}\right)}\frac{1}{r^3} + \dots, \\ &p(r) &= \ -\frac{i\sqrt{\frac{2}{3}}\left(s^2+3\sqrt{1-s^2}-1\right)}{s^3}\frac{1}{r^2} + \dots, \\ &q(r) &= \ -\frac{3i\left(\sqrt{6}\sqrt{1-s^2}\right)}{s}r + \frac{\sqrt{\frac{2}{3}}i\sqrt{1-s^2}\left(5s^2+9\sqrt{1-s^2}-5\right)}{3s^3}\frac{1}{r} + \dots, \\ &f(r) &= \ \frac{1-s^2+\sqrt{1-s^2}}{s^2} + \frac{2\left(-2+2s^2-(2+s^2)\sqrt{1-s^2}\right)}{9s^4}\frac{1}{r^2} + \frac{\kappa}{r^3} + \dots. \end{split}$$

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The parameter κ is uniquely determined by requiring this to extend to a smooth solution on the ball $M_6 \cong B_6$. As an expansion in

$$\delta = \sqrt{-1 + \mathrm{s}^{-1}}$$

this is

$$\frac{3\sqrt{3}}{4}\kappa = \delta + \frac{\sqrt{2}}{3}\delta^2 + \frac{113}{36}\delta^3 + \frac{25}{9\sqrt{2}}\delta^4 + \frac{1127}{288}\delta^5 + \frac{35}{9\sqrt{2}}\delta^6 + \dots$$

Similar results hold in the 1/4 BPS case, except here we find a *two-parameter* family of solutions, leading to a new supersymmetric squashing of S^5 . In particular this includes a one-parameter subfamily of 1/2 BPS solutions.

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As in four dimensions the regularized action is

$$-\log Z_{\text{gravity}} = S_{\text{bulk}} + S_{\text{Gibbons}-\text{Hawking}} + S_{\text{ct}}$$
.

However, unlike in four dimensions the counterterms S_{ct} had not been computed.

This is a straightforward, but very long, computation. In particular the **B**-field is both massive and has a cubic Chern-Simons interaction, which leads to a much more complicated analysis than for more standard fields.

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$$\begin{split} S_{ct} &= \frac{1}{8\pi G_N} \int_{\partial M_0} \left\{ \Big[\frac{4\sqrt{2}}{3} + \frac{1}{2\sqrt{2}} R(h) - \frac{1}{6\sqrt{2}} \|B\|_h^2 + \frac{3}{4\sqrt{2}} R(h)_{ij} R(h)^{ij} - \frac{15}{64\sqrt{2}} R(h)^2 - \frac{3}{4\sqrt{2}} \|F_I\|_h^2 \right. \\ &+ \frac{1}{12\sqrt{2}} \mathrm{Tr}_h B^4 + \frac{5}{8\sqrt{2}} \|d \ast_h B + \frac{i\sqrt{2}}{3} B \wedge B\|_h^2 - \frac{1}{4\sqrt{2}} \langle B, \mathrm{d}\delta_h B + \frac{i\sqrt{2}}{3} \mathrm{d} \ast_h B \wedge B \rangle_h - \frac{1}{\sqrt{2}} \|\mathrm{d}B\|_h^2 \\ &+ \frac{4\sqrt{2}}{3} (1-X)^2 - \frac{1}{\sqrt{2}} \langle \mathrm{Ric}(h) \circ B, B \rangle_h + \frac{9}{32\sqrt{2}} R(h) \|B\|_h^2 - \frac{13}{192\sqrt{2}} \|B\|_h^4 \Big] \sqrt{\det h} \, \mathrm{d}^5 x \\ &- \frac{1}{4\sqrt{2}} B \wedge \left[\mathrm{d} \ast_h \mathrm{d}B + \frac{\sqrt{2}i}{3} B \wedge \delta_h B - \frac{2}{9} B \wedge \ast_h (B \wedge B) \right] \Big\} \,. \end{split}$$

Here $\operatorname{Ric}(\mathbf{h})_{ij} = \mathbf{R}(\mathbf{h})_{ij}$ denotes the Ricci tensor of the boundary metric \mathbf{h}_{ij} , with $\mathbf{R}(\mathbf{h})$ the Ricci scalar. The inner product of two **p**-forms ν_1 , ν_2 is defined by $\langle \nu_1, \nu_2 \rangle_h \sqrt{\det h} d^5 \mathbf{x} = \nu_1 \wedge *_h \nu_2$, which then also defines the square norm via $\|\nu\|_h^2 = \langle \nu, \nu \rangle_h$. The adjoint δ_h of **d** with respect to \mathbf{h}_{ij} acting on the two-form **B** is $\delta_h \mathbf{B} = *_h \mathbf{d} *_h \mathbf{B}$, and we have also defined $\operatorname{Tr}_h \mathbf{B}^4 \equiv \mathbf{B}_i{}^j \mathbf{B}_j{}^k \mathbf{B}_k{}^l \mathbf{B}_l{}^i$. Finally, we have defined the **p**-form $(\mathbf{S} \circ \nu)_{i_1 \cdots i_p} \equiv \mathbf{S}_{[i_1}{}^j \nu_{|j|i_2 \cdots i_p]}$, where \mathbf{S}_{ij} is any symmetric 2-tensor, and ν is any **p**-form.

Using this we may compute the holographic free energy. For example, for the 3/4 BPS solution we find

$$\begin{split} \mathsf{S}_{\text{bulk}} + \mathsf{S}_{\text{Gibbons-Hawking}} + \mathsf{S}_{\text{ct}} &= -\frac{27\pi^2}{4\mathsf{G}_{\mathsf{N}}} \left(1 + \frac{8}{3}\delta^2 + \frac{16\sqrt{2}}{27}\delta^3 + \frac{68}{27}\delta^4 \right. \\ &+ \frac{28\sqrt{2}}{27}\delta^5 + \frac{32}{27}\delta^6 + \dots \right) \,. \end{split}$$

This agrees with the field theory result. The BPS Wilson loop maps to a fundamental string in type **IIA**, at the "pole" of the internal S^4 [Assel-Estes-Yamazaki]. The renormalized string action is

$$S_{\rm string} = \int_{\Sigma} \left[X^{-2} \sqrt{\det \gamma} \, d^2 x + i B \right] - \frac{3}{\sqrt{2}} {\rm length}(\partial \Sigma) ,$$

and also agrees with the large \mathbf{N} field theory results.

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